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Liverpool Telescope classification of optical transients

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 on **22 Oct 2018; 22:36 UT**
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We conducted spectroscopic observations of seven Local Group nova candidates using the SPRAT instrument ([Pascik et al. 2014](#)) on the 2m-robotic Liverpool Telescope (LT; [Steele et al. 2004](#)). Each spectral observation is preceded by a series of (typically at least 3) 10s acquisition images taken either by SPRAT (unfiltered) or **IO:O** through a Sloan-r' filter. Each spectrum consists of 3×600s or 3×900s exposure time using the blue-optimised mode of SPRAT.

We attempted to obtain a spectrum of [AT2018hfs](#) on 2018-10-15.97 UT, but the object is barely visible in the acquisition images. [AT2018ggg](#) and [AT2018had](#) were observed at 2018-09-22.95 and 2018-10-04.12 UT respectively. A faint continuum is detected in each case, but the S/N is too low for a classification. [AT2018hho](#) was observed at 2018-10-16.94 UT. We detect a H α emission line with FWHM ~ 3100 km/s. A spectrum of [AT2018hcd](#) taken 2018-10-16.98 UT shows a red continuum. The S/N is low, but there is tentative evidence for TiO absorption bands, indicating the transient is likely a red LPV.

We obtained a spectrum of [AT2018hhy](#) at 2018-10-17.99 UT, which shows TiO absorption bands on a red continuum, indicating that this object is a red LPV. We also note that there is a variable coincident with the position of [AT2018hhy](#) in [Macri et al. \(2001\)](#) and [Hartman et al. \(2006\)](#).

We obtained a spectrum of [AT2018hld](#) at 2018-10-18.89 UT. The spectrum shows strong Balmer emission and we measure the H α FWHM ~ 2300 km/s. The spectrum also shows numerous Fe II emission lines, including those of the 42, 48, 49 and 74 multiplets, along with O I 7774 Å (the latter of which seems to show a P-Cygni profile). The spectrum confirms [AT2018hld](#) is a nova in M31 and a member of the Fe II spectroscopic class. The spectrum is also broadly similar with that obtained around one day earlier by Robin Leadbeater (see [AT2018hld TNS page](#)).

The objects that we were able to classify are summarised below:

Nova candidate	Date (UT)	Classification

AT2018hcd	2018-10-16.98	Red LPV
AT2018hhy	2018-10-17.99	Red LPV
AT2018hld	2018-10-18.89	Fe II Nova

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